

Weak measurements and metrology

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Collaborators



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Weak measurements

Discovered during exotic research on foundations of QM

Aharonov-Albert-Vaidman PRL 1988



➤ **Explore foundations of QM**

Hardy paradox Aharonov et al. PLA 2002

Observe wavefunction Lundeen et al. Nature 2011, Steinberg et al. Science 2011

➤ **Cavity QED experiments** Wiseman PRA 2002

➤ **Telecom optics** Brunner et al. PRL 2003

➤ **Superluminal propagation** Brunner et al. PRL 2004

➤ **Solid-state physics** Williams-Jordan PRL 2008

WM and Metrology



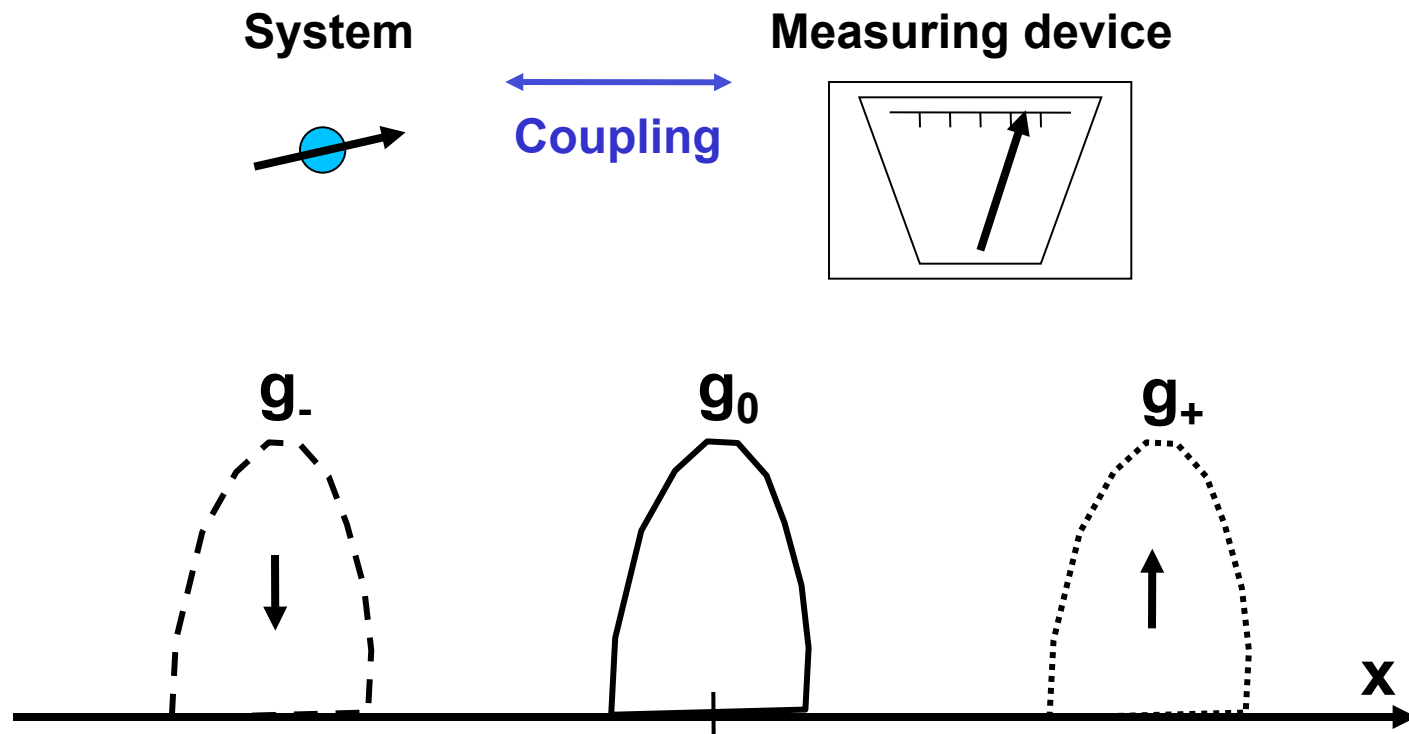
QUANTUM INSPIRED METROLOGY

- **Spin Hall effect of light** (Hosten & Kwiat, Science 2008)
- **Small beam deflections ~15fm** (Dixon, Starling, Jordan, Howell PRL 2009)

HERE : SMALL PHASE SHIFTS, MAGNETOMETRY

What is a weak measurement?

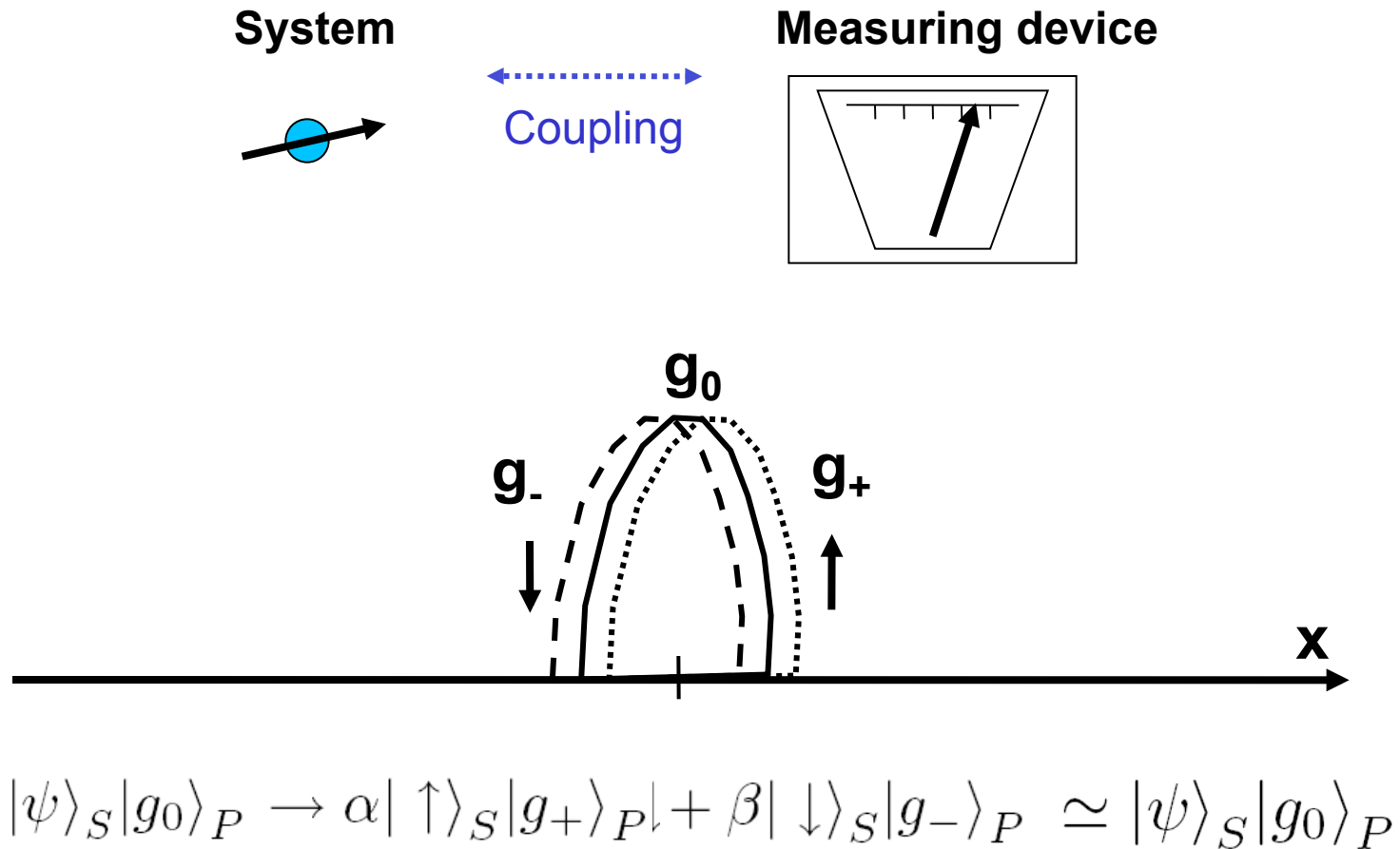
Standard picture of Q measurement (von Neumann)



$$|\psi\rangle_S |g_0\rangle_P = (\alpha |\uparrow\rangle_S + \beta |\downarrow\rangle_S) |g_0\rangle_P \rightarrow \alpha |\uparrow\rangle_S |g_+\rangle_P + \beta |\downarrow\rangle_S |g_-\rangle_P$$

Pointer read-out \rightarrow Collapse of system / Full information

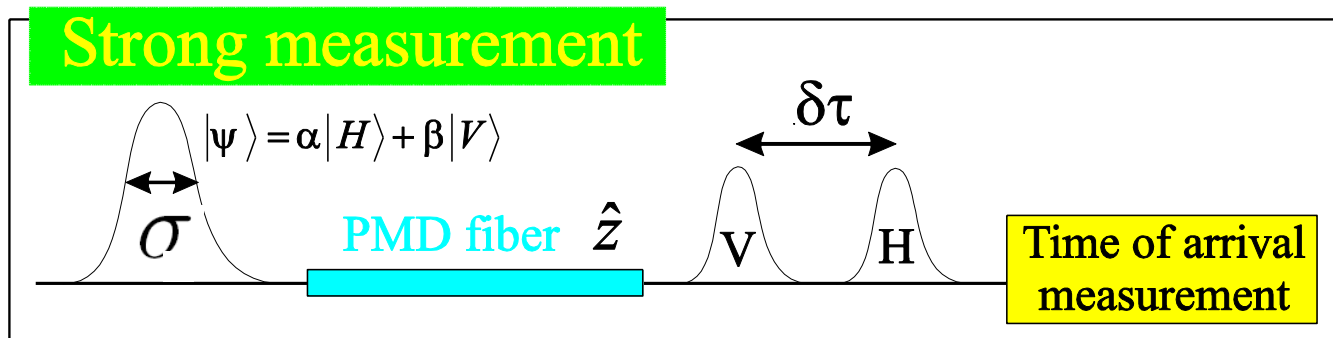
What is a weak measurement?



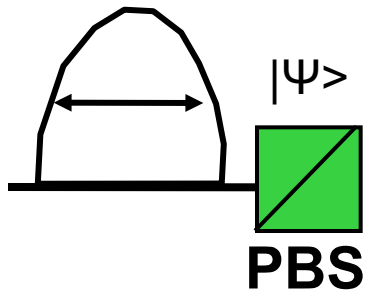
Pointer read-out \rightarrow no disturbance / little information

Example - Birefringence

Coupling between polarization and temporal mode

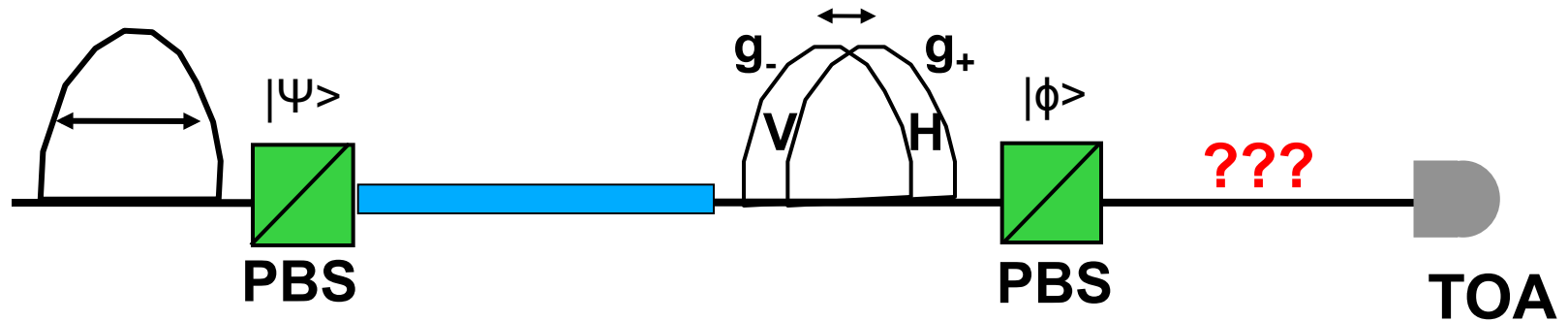


Post-selection - Weak value



1. Pre-selection $|\psi\rangle = \alpha|H\rangle + \beta|V\rangle$

Post-selection - Weak value

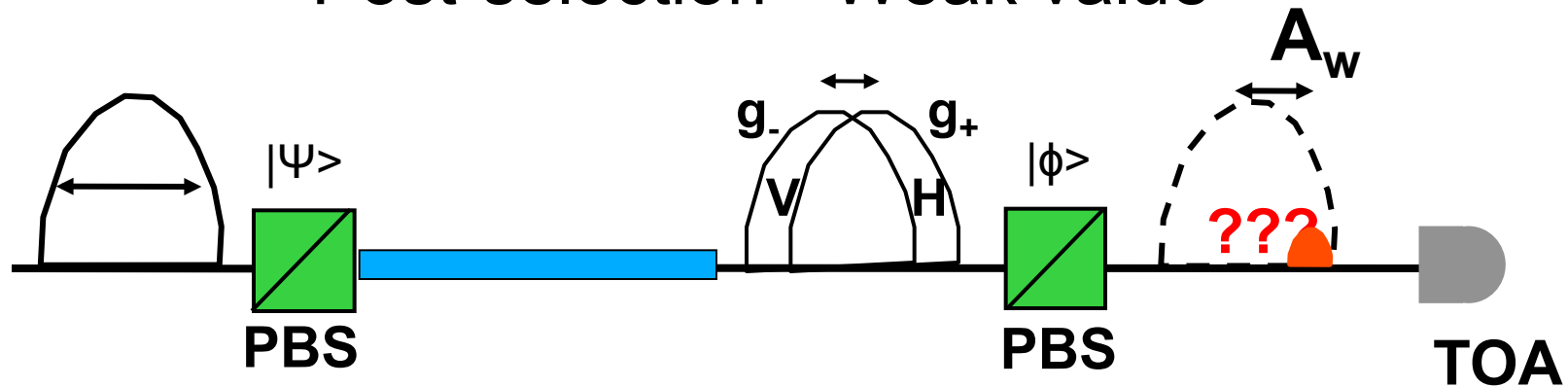


1. Pre-selection $|\psi\rangle = \alpha|H\rangle + \beta|V\rangle$

2. Weak meas. $U = e^{-i\tau\omega\sigma_z}$ Shift operator

3. Post-selection $|\phi\rangle = \mu|H\rangle + \nu|V\rangle$

Post-selection - Weak value



1. Pre-selection $|\psi\rangle = \alpha|H\rangle + \beta|V\rangle$

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Output state

$$\begin{aligned} |g_{out}\rangle &= \langle\phi|e^{-i\tau\omega\sigma_z}|g(t)\rangle|\psi\rangle \\ &\simeq \langle\phi|\psi\rangle e^{-i\tau A_w\omega}|g(t)\rangle \end{aligned}$$

Amplification of pointer shift

$$A_w = \frac{\langle\phi|\sigma_z|\psi\rangle}{\langle\phi|\psi\rangle}$$

Weak value

Fastlight & Neutrinos

Desktop experiment

Brunner, Scarani, Legre, Wegmuller, Gisin, PRL 2004

NEUTRINOS OSCILLATIONS \approx BIREFRINGENCE

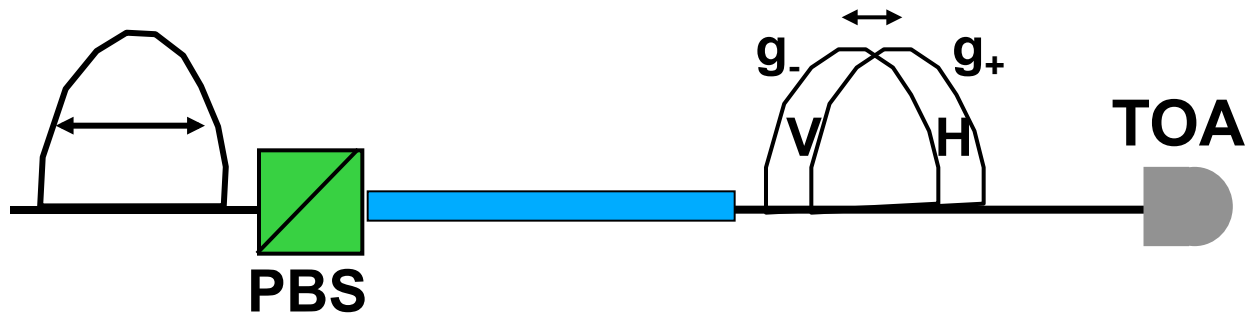
SUPERLUMINAL NEUTRINOS AT OPERA ?

PROBABLY NOT

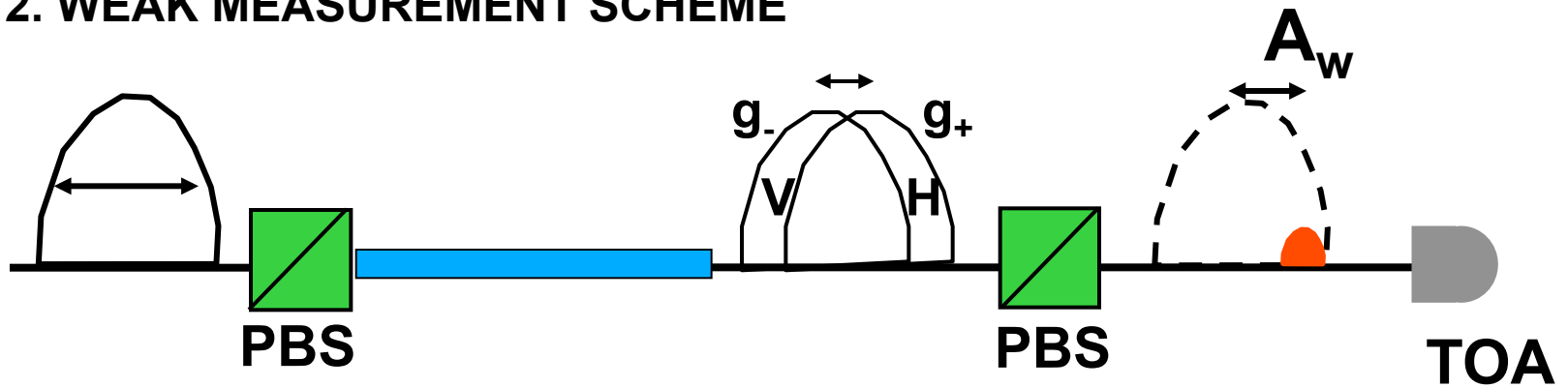
Berry, Brunner, Popescu, Shukla, J Phys A 2011

Metrology

1. DIRECT OBSERVATION



2. WEAK MEASUREMENT SCHEME



WM scheme

Advantage: Signal is amplified by the weak value

$$\text{Amplification} \quad A_w = \frac{\langle \phi | A | \psi \rangle}{\langle \phi | \psi \rangle}$$

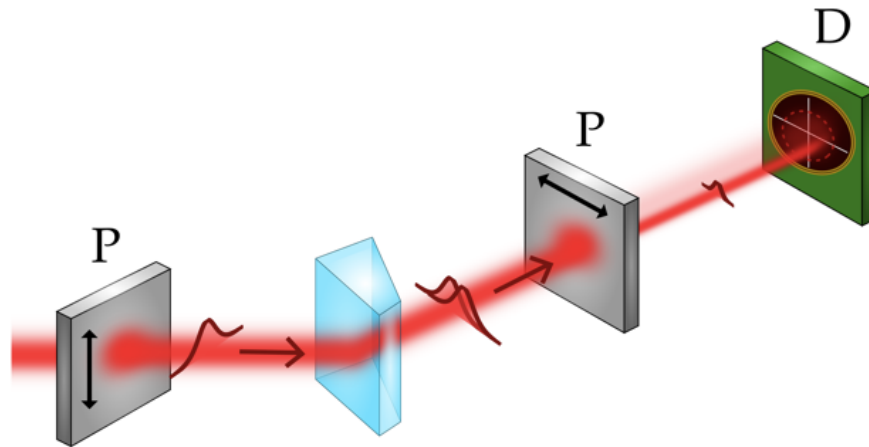
Price to pay: intensity of the signal is reduced

$$\text{Prob. success} \quad p \simeq \frac{1}{A_w^2}$$

Useful for transverse effects

1. Spin Hall effect of light (Hosten-Kwiat 2008)

Effect $\sim \text{\AA}$



2. Small beam displacement (Dixon et al. PRL 2009)

Effect $\sim 15\text{fm}$



What about longitudinal effects?

Measuring small phase shifts



Small phase shifts stable in time

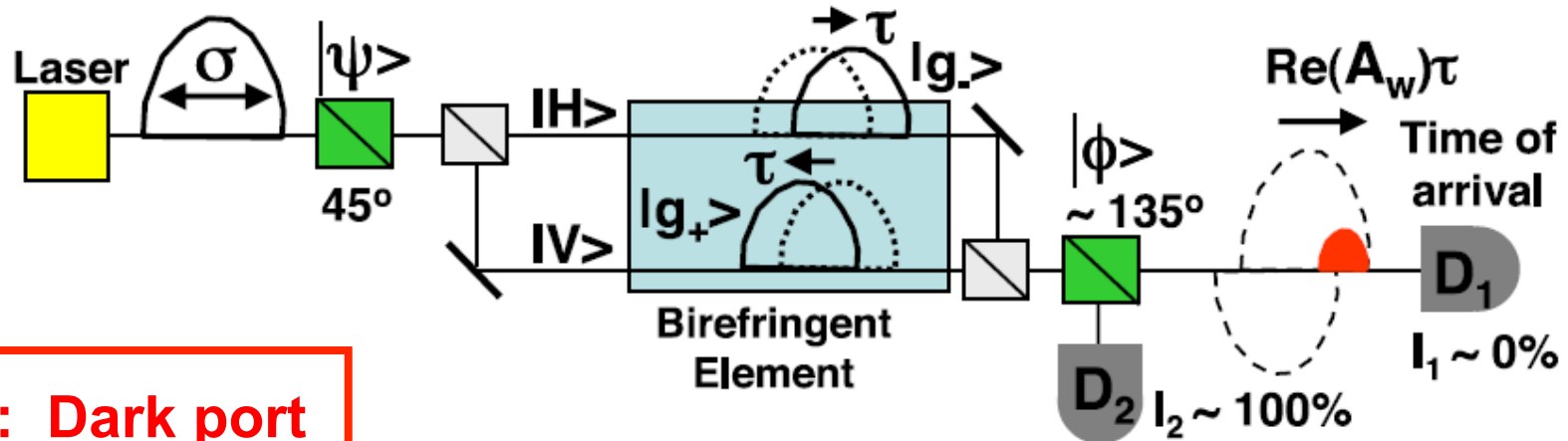
Limiting factor: alignment errors (not photon statistics)

Compare 3 techniques

- a) **Weak meas. Real weak value**
- b) **Weak meas. Imaginary weak value**
- c) **Standard interferometry**

Different operating point of IF and detection

a) Weak Meas. Real weak value



Time resolution
of detector

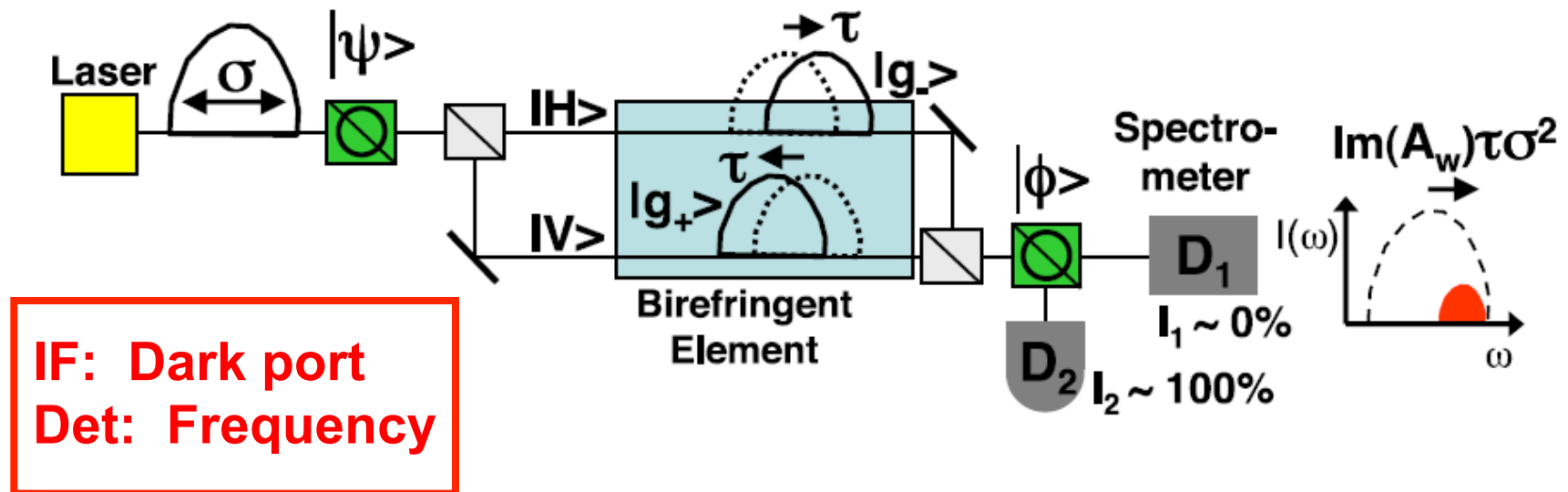
Signal: $\tau A_w \simeq \tau / \sqrt{p} > \Delta t \longrightarrow \tau > \Delta t \sqrt{p}$

Angle misalignment ϵ : $P_{error} \simeq \epsilon^2 \longrightarrow p > \epsilon^2$

Resolution limit:

$$\tau > \epsilon \Delta t$$

b) Weak Meas. Imaginary weak value



Signal: $\delta\omega \simeq \tau / (\sigma^2 \sqrt{p}) \longrightarrow \delta\omega > \Delta\omega$

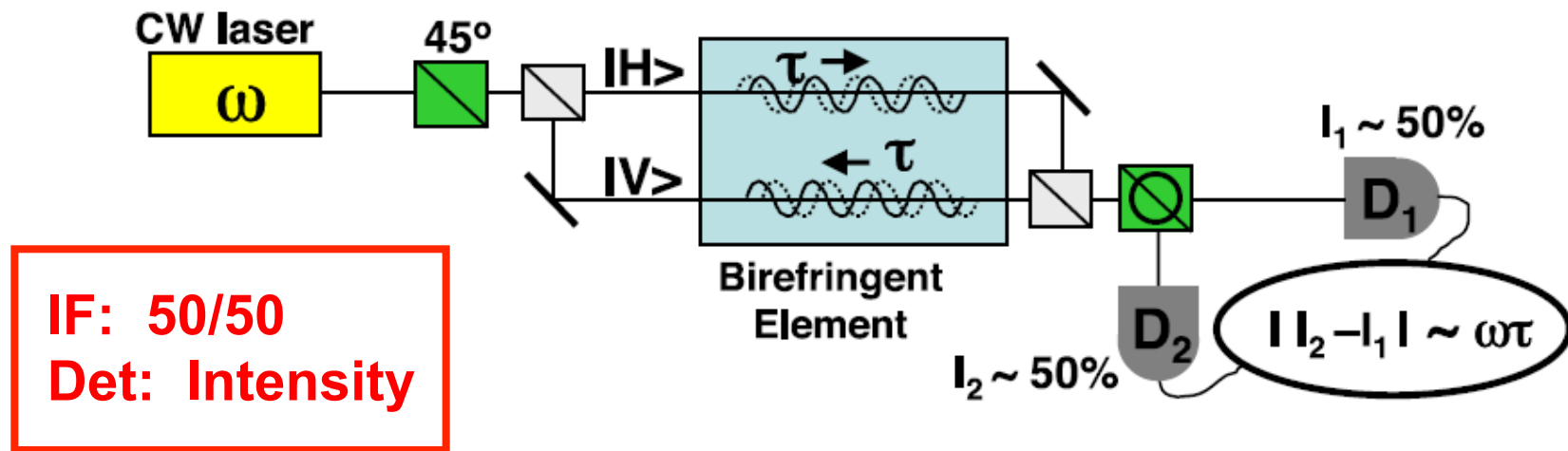
Spectral resolution
↓

Angle misalignment ϵ : $P_{error} \simeq \epsilon^2 \longrightarrow p > \epsilon^2$

Resolution limit:

$$\tau > \epsilon \sigma^2 \Delta\omega$$

c) Standard Interferometry



Signal: $|I_1 - I_2| \simeq 2N\omega\tau$

Angle misalignment ϵ : $P_{error} \simeq \epsilon \longrightarrow \omega\tau > \epsilon$

Resolution limit:

$$\tau > \frac{\epsilon}{\omega}$$

Comparison

Resolution limit

In practice

a) Real weak value

$$\tau > \epsilon \Delta t$$

$$\Delta t > 10 \text{ ps}$$

Gol'tsman et al. (2005)

c) Interferometry

$$\tau > \frac{\epsilon}{\omega}$$

$$1/\omega \sim 1 \text{ fs}$$

Optical frequencies

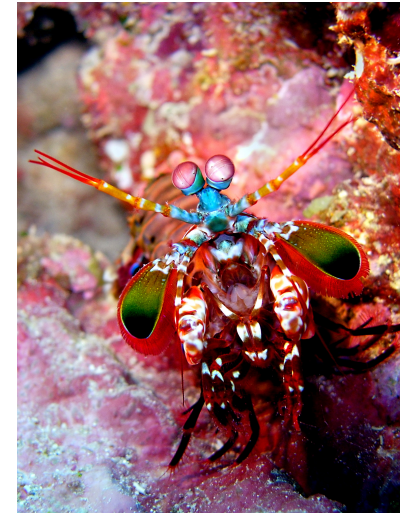
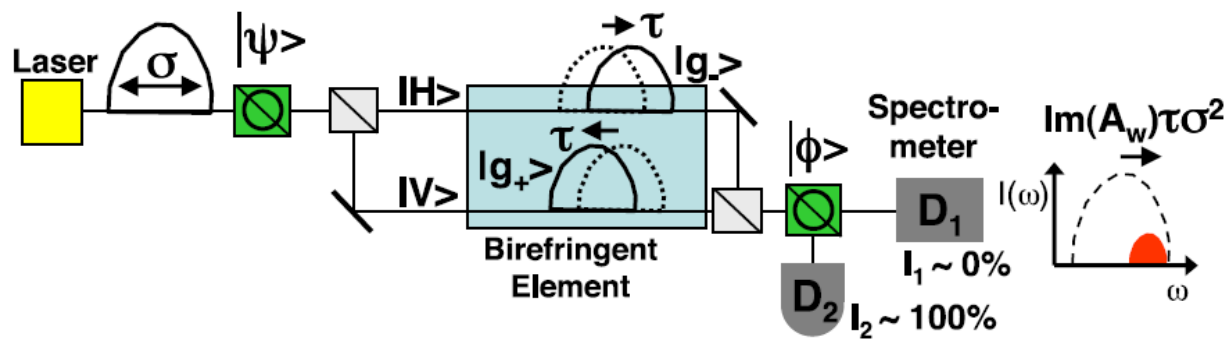
b) Imag. weak value

$$\tau > \epsilon \sigma^2 \Delta \omega / 2$$

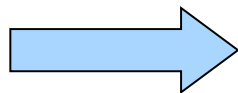
$$\sigma^2 \Delta \omega = 0.5 \text{ as}$$

Perspectives

PRACTICAL IMPLEMENTATION ?



Roberts et al. Nature Phot 2009



OTHER METROLOGY APPLICATIONS ?
NASA ? LIGO ?

More

**Weak measurements
amplification**



**Noiseless amplification
of coherent states**

Ralph-Lund 2009

Exp : Pryde, Bellini, Leuchs

- Based on post-selection
- Heralded
- Achieve something that would be impossible deterministically

Brunner, Polzik, Simon PRA 2011

Heralded amplification of rotations in spin ensembles



MAGNETOMETRY

Thank you